

Radiation of ultra-relativistic electrons and positrons during channeling in crystals and in the fields of powerful lasers



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Channelling: CERN NA63-Collaboration (Aarhus, U.I.Uggerhoj et.al.

Laser: 1. Di Piazza, A., Keitel C.H., et.al
Max-Planck-Institut für Kernphysik, Heidelberg, Germany

2. Russia: Kostjukov I.Yu. et.al. Institute of Applied Physics of the RAS, Nizhny Novgorod

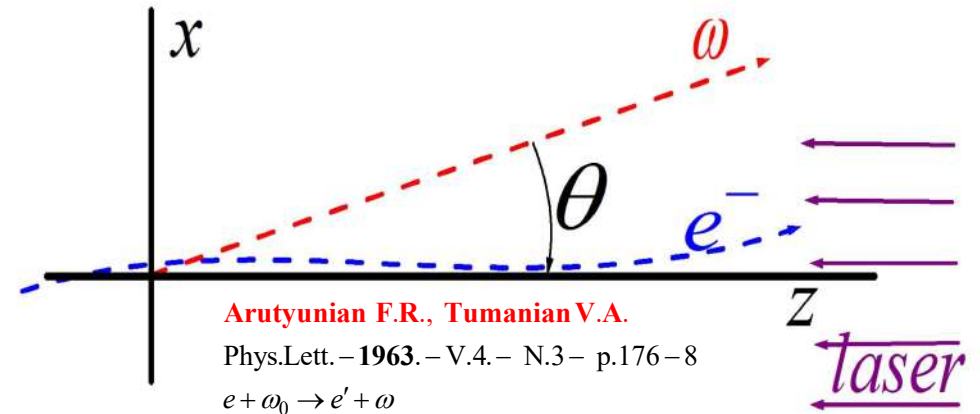
$$\mathbf{E} = -\mathbf{E}_0 \sin(\delta), \quad (1) \quad \gamma = \gamma_0 \sqrt{1 + v_0^2}$$

$$\delta = \omega_0(t + z/c),$$

$$x_m = \begin{cases} \sqrt{2} \frac{v_0 c}{\Omega_0 \gamma}, & \text{laser} \\ \frac{d_p}{2} \sqrt{\frac{E_{\perp}}{U_m}}, & \text{crystal} \end{cases} \quad (2)$$

$$\Omega_0 = \begin{cases} (1 + \beta_0) \omega_0 \approx 2\omega_0, & \text{laser} \\ (2c / d_p) \theta_L, & \text{crystal}, \end{cases} \quad (3)$$

• **Laser** $e + k\omega_0 \rightarrow e' + \omega \quad (4)$



Milburn R.N. Phys.Rev.Lett. – 1963. – V.10. – N.3. – P.75–77

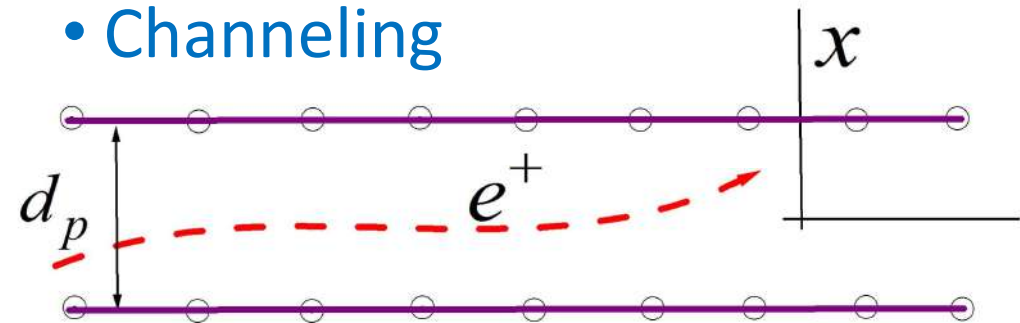
Nikishov A.I., Ritus V.I. Sov.Phys.JETP. 1964. V.19. P.1191

Brown L.S., Kibble T.W.B. Phys.Rev. 1964. V. 133. P.705

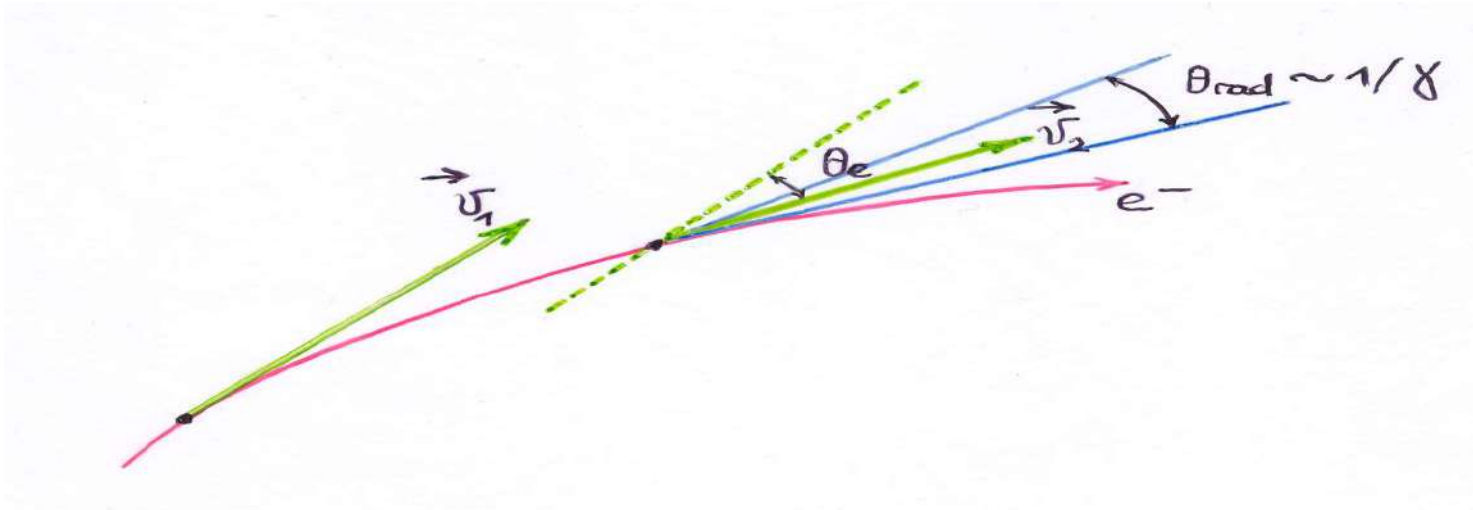
Bula C., McDonald K.T., Prebys E.J., et.al. // Phys. Rev. Lett. 1996. V.76. P.3116.

Khokonov M.Kh., Carrigan R.A. Fermilab preprint, Pub-97/115, 1997; NIMB, V.145, P.133-141, (1998);

• **Channeling**



Radiation by relativistic electron



- **Non-dipole parameter**

$$D \approx \theta_e / \theta_{\text{rad}} \approx \theta_e \gamma \quad (1)$$

- $D \gg 1$ - Synchrotron approximation $D^2 = \overline{\beta_{\perp}^2} \gamma^2 \quad (2)$

- $D \ll 1$ - Dipole approximation $\beta_{\perp} \gamma$ - Lorenz-invariant

Invariants

$$v_0^2 = \frac{e^2 \mathcal{E}_0^2}{2m^2 c^2 \omega_0^2} = \frac{e^2}{m^2 c^4} A^2 \quad (1)$$

$$a = \frac{2\hbar}{m^2 c^2} kp = \frac{2\hbar \Omega_0 \gamma}{m c^2} \quad (2)$$

$$\chi = \frac{e\hbar |F_{\mu\nu} p^\nu|}{m^2 c^3} \approx \frac{2e\mathcal{E}_0 \hbar}{m^2 c^3} \gamma \quad (3)$$

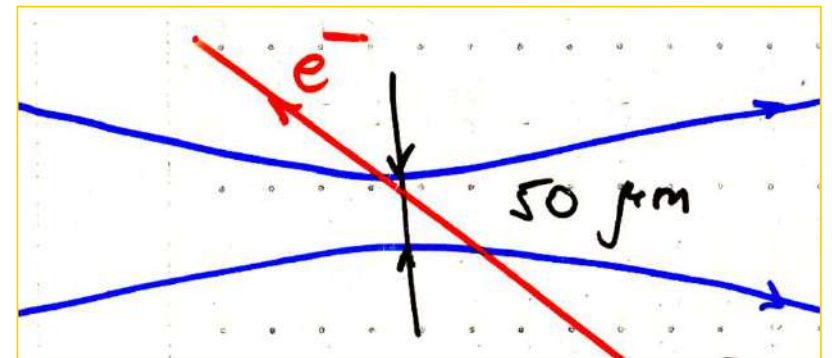
$$\chi \approx a v_0$$

$\Omega_0 = 2\omega_0$ – frequency of transverse motion

$$\bar{\gamma}' = \sqrt{1 + v_0^2}$$

in the average

rest frame - ARF



Strong field effects

$$\chi = \frac{\mathcal{E}}{\mathcal{E}_c} \gamma \geq 1$$

$$\mathcal{E}_c = \frac{m^2 c^3}{e\hbar} \rightarrow I \approx 10^{28} \frac{\text{W}}{\text{cm}^2}, \quad F \approx 10^{16} \frac{\text{eV}}{\text{cm}}$$

crystals $F \approx \frac{2Ze^2}{da_F} \sim 10^3 \frac{\text{eV}}{\text{Å}} = 10^{11} \frac{\text{eV}}{\text{cm}}$ lab frame

$$F' = \gamma F \approx 10^{16} \frac{\text{eV}}{\text{cm}} \quad \text{rest frame for } \gamma \approx 10^5$$

Khokonov M.Kh., Nitta H., *Phys. Rev. Lett.*, V.89, 094801 (2002)

Nitta H., Khokonov M.Kh., Nagata Y., Onuki S. *Phys. Rev. Lett.*,
V. 93, 180407 (2004)

Comparison – lasers, crystals

Crystal (axial):

$$E_{crystal}^2 \approx K \frac{\pi Z^2 e^2}{d} N$$

$$520 \text{ eV/\AA}^3 \text{ Si } \langle 110 \rangle$$

$$580 \text{ eV/\AA}^3 \text{ C } \langle 100 \rangle$$

$$1000 \text{ eV/\AA}^3 \text{ W } \langle 111 \rangle$$

Laser:

10 TW,

beam radius of 50 μm

$$E_{laser}^2 \approx 700 \text{ eV/\AA}^3$$

$$v_0 = 0.3$$

20-200 MeV electrons

For 1 TW laser, $\lambda = 1 \mu\text{m}$

$$\hbar\omega \sim 6 - 600 \text{ keV}$$

Channeling radiation:

$$\hbar\omega \sim 6 - 200 \text{ keV}$$

$$D = \langle \beta_{\perp}^2 \gamma^2 \rangle^{1/2} = \frac{1}{\sqrt{2}} \frac{x_m \Omega_0}{c} \gamma = \begin{cases} v_0, \\ \sqrt{E_{\perp} E} / (mc^2), \end{cases}$$

для лазера,

для кристалла,

Equations of motion

$$\frac{d\mathbf{p}}{dt} = e\mathcal{E} + e\boldsymbol{\beta} \times \mathbf{H} \quad (1)$$

$$x = c\alpha(\Omega_0\gamma_0)^{-1} \sin \delta$$

$$z = \beta_0 ct - c\alpha^2(8\Omega_0\gamma_0^2)^{-1} \sin 2\delta, \quad (2)$$

$$t = \Omega_0^{-1} \delta + \alpha^2(8\Omega_0)^{-1} (1 - \beta_0) \sin 2\delta$$

$$\gamma_0 = (1 - \beta_0^2)^{-1/2} \quad (6)$$

$$\alpha^2 = 2v_0^2 / (1 + v_0^2)$$

$$\boxed{\beta_{\perp}^2 \gamma^2 = v_0^2} \quad (7)$$

$$\gamma(t) = \gamma_0 \sqrt{1 + v_0^2} \left[1 + \frac{v_0^2}{2(1 + v_0^2)\gamma_0^2(1 + \beta_0)} \cos(2\delta) \right] \quad (3)$$

if $\gamma_0 \gg 1$, and $\gamma_0 \gg v_0$, then $\gamma = \gamma_0 \sqrt{1 + v_0^2}$ (4)

$$\delta \approx \Omega_0 t \quad (5)$$

$$\boxed{\begin{aligned} x(t) &= x_m \sin(\Omega_0 t), \\ z(t) &= ct \left(1 - \frac{1}{2\gamma^2} - \frac{\langle \beta_{\perp}^2 \rangle}{2} \right) - \frac{\Omega_0}{8c} x_m^2 \sin(2\Omega_0 t), \end{aligned}} \quad (8)$$

Radiation spectrum

$$\frac{d^3 N}{dud(\Omega_0 t)d\varphi} = \frac{2\alpha D^2}{\pi a} \sum_{k=k_m}^{\infty} \left[g_k + \frac{uu'}{2} \left(g_k + \frac{i_z^2}{2D^2} \right) \right], \quad (1) \quad u = \hbar\omega / E$$

$$g_k = i_x^2 + \frac{1}{2D^2} \eta_k^2 i_z^2 - \sqrt{2} \frac{\eta_k}{D} i_x i_z \cos\varphi, \quad (2) \quad k_m = 1 + \text{Int} \left(\frac{1 + D^2}{a} u' \right) \quad (6) \quad u' \equiv u / (1 - u) \quad (7)$$

$$\eta_k = \left(\frac{ak}{u} - ak - D^2 - 1 \right)^{1/2}, \quad (3)$$

$$0 < u < u_{km} = \frac{ak}{1 + D^2 + ak}, \quad (8)$$

$$i_x = \frac{1}{A} (kS_1 + S_2), \quad i_z = \frac{1}{B} S_2, \quad (4)$$

$$S_1 = \sum_{m=-\infty}^{\infty} J_m(B) J_{k+2m}(A), \quad A = \frac{2\sqrt{2}}{a} u' D \eta_k \cos\varphi, \quad (5)$$

$$S_2 = \sum_{m=-\infty}^{\infty} m J_m(B) J_{k+2m}(A), \quad B = \frac{u'}{2a} D^2$$

1. Dipole approximation $D \ll 1$

$$\frac{1}{n_0} \frac{dN}{du} = \frac{3}{2a} \left[1 - \frac{2u}{a(1-u)} + \frac{2u^2}{a^2(1-u)^2} + \frac{u^2}{2(1-u)} \right], \quad u \leq u_m \quad (1) \quad u_m = \frac{\omega_m}{E} = \frac{a}{1+a} \quad (2) \quad \frac{dN}{d(\Omega_0 t)} = \frac{2}{3} \alpha D^2 \equiv n_0 \quad (3)$$

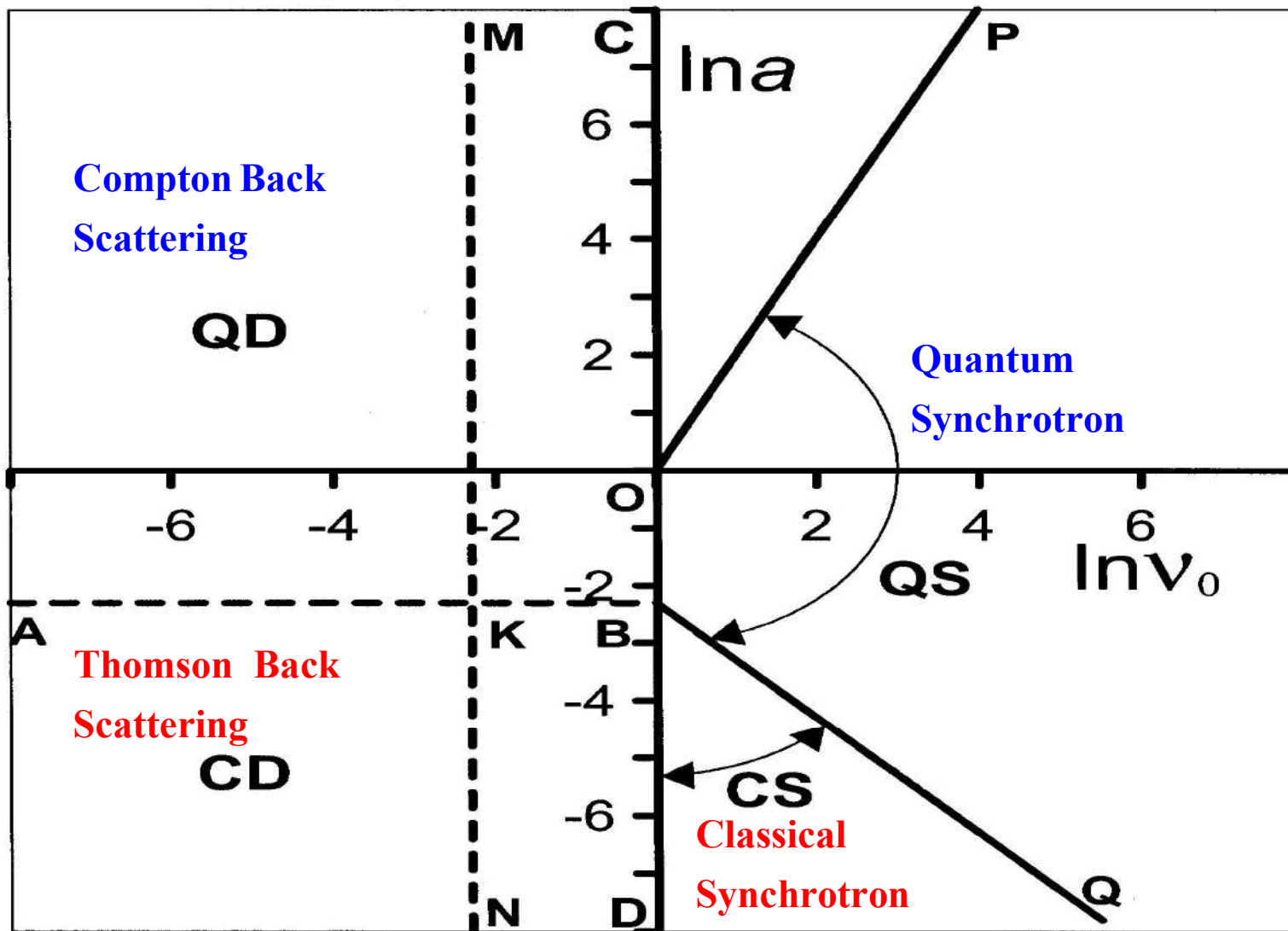
2. Constant field approximation - CFA $\frac{D^2}{a} > 1$, $\left(\text{or } \frac{D^2}{a} \frac{u}{1-u} > 1 \right)$ - CFA (4)

$$\frac{d^2 N(t)}{dud(\Omega_0 t)} = \frac{2}{3} \alpha \frac{\sqrt{3}}{\pi a} J_{SYN}(\xi) \quad (5) \quad \xi = \frac{2}{3\chi} \frac{u}{1-u} \quad (6) \quad \chi = aD \sin \alpha_0 / \sqrt{2} \quad (7)$$

$$J_{SYN}(\xi) = \left(2 + \frac{u^2}{1-u} \right) K_{2/3}(\xi) - \int_{\xi}^{\infty} K_{1/3}(x) dx = \quad (8)$$

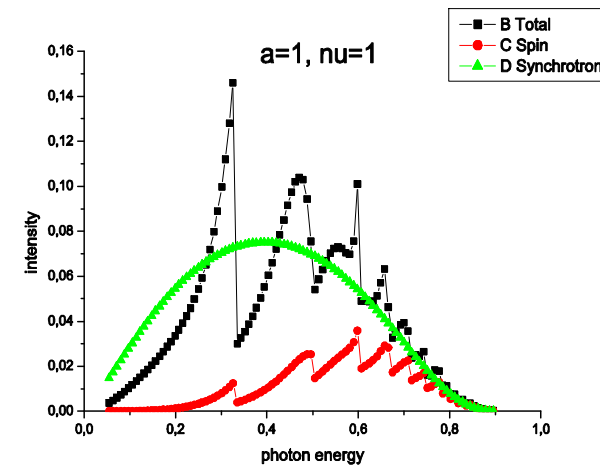
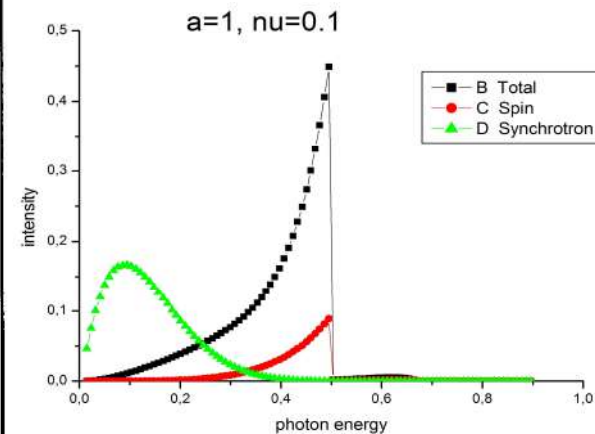
$$= \sqrt{3} \int_0^{\infty} \left[\frac{3 + 12x^2 + (16/3)x^4}{3 + 4x^2} + \frac{u^2}{1-u} \left(1 + \frac{2}{3}x^2 \right) \right] \frac{\exp \left[-\xi \left(1 + 4x^2 / 3 \right) \sqrt{1 + x^2 / 3} \right]}{\sqrt{1 + x^2 / 3}} dx \quad (9)$$

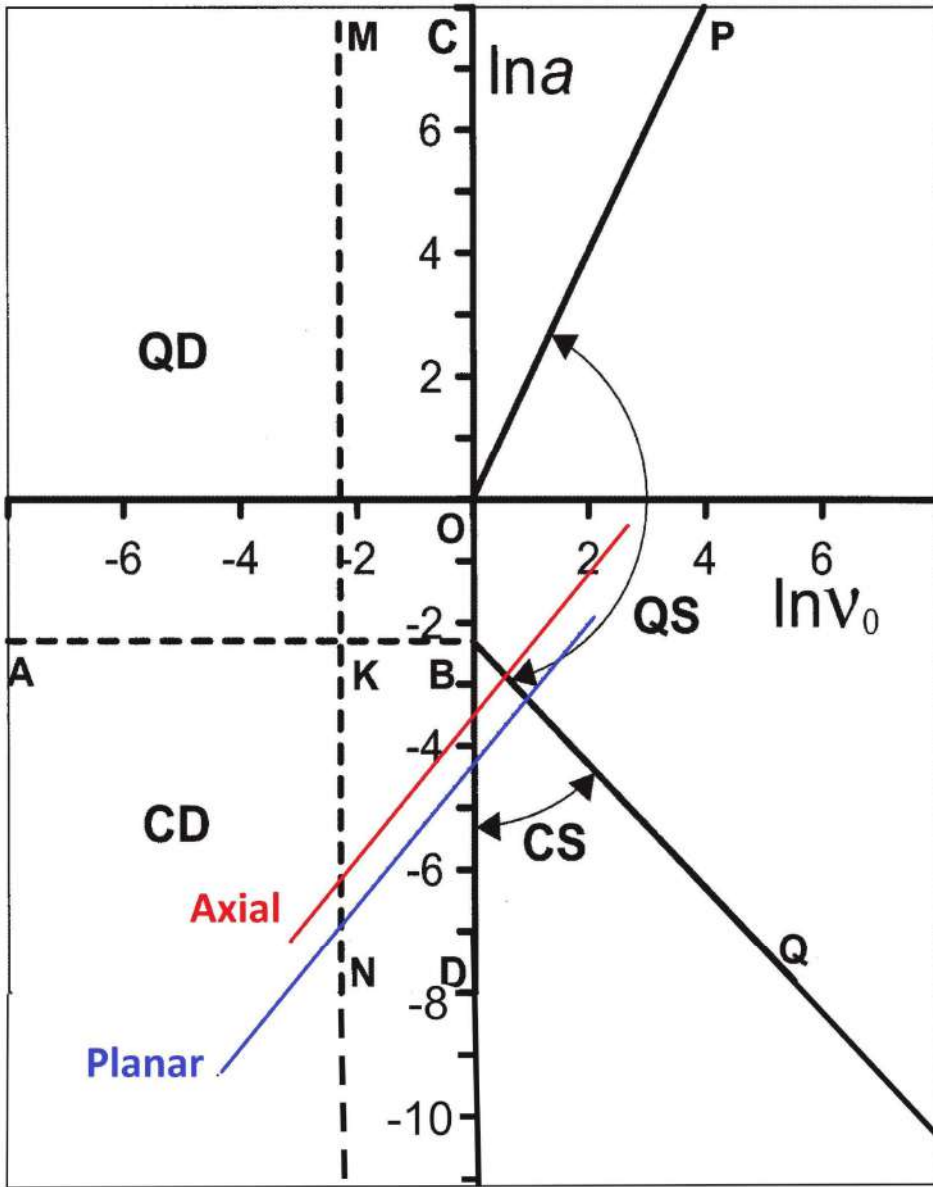
General Classification



$E = 125 \text{ GeV}$

$\lambda = 1 \mu\text{m}$





Channeling

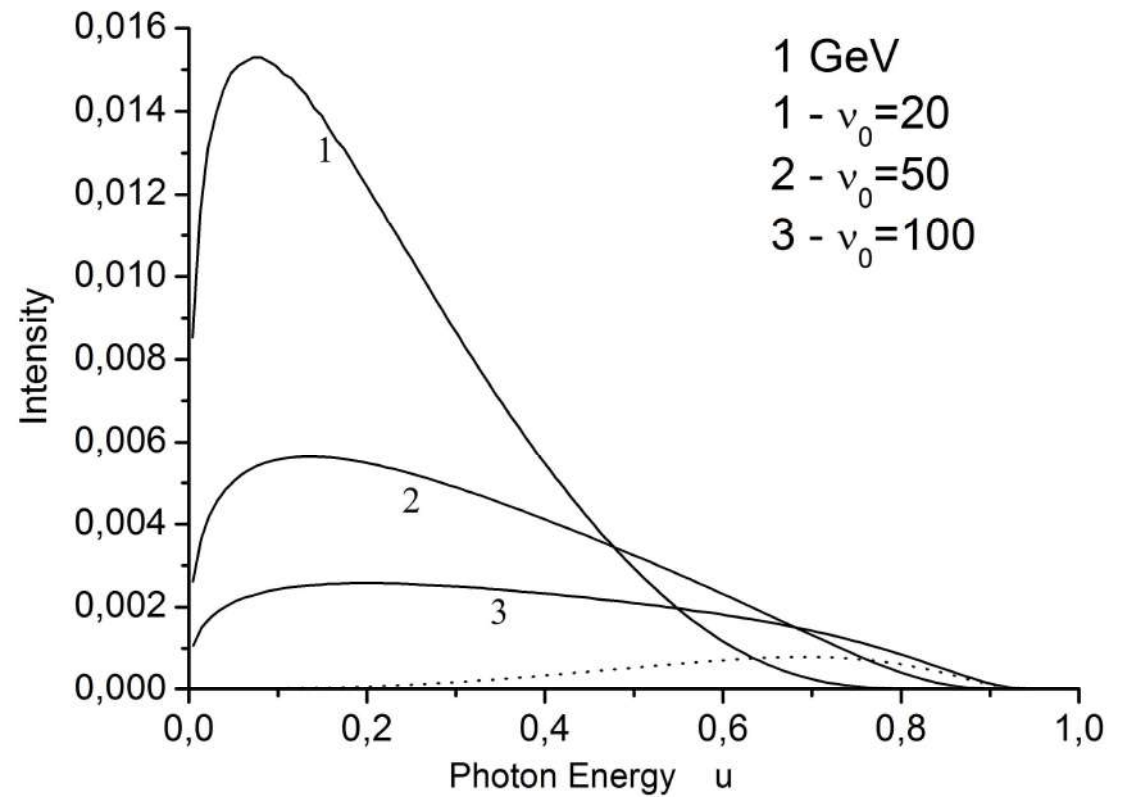
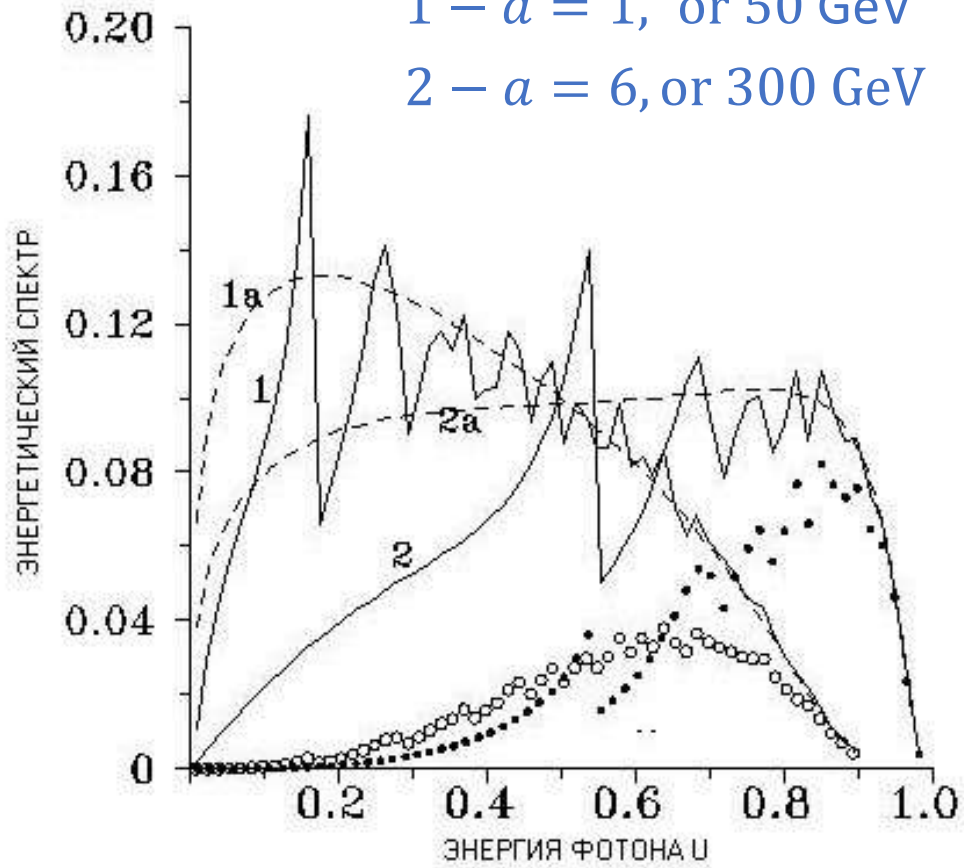
$$D = a \frac{mc^2 d_p}{4\hbar c} \sqrt{\frac{E_{\perp}}{2U_m}} \quad (1)$$

$$a \approx \frac{2\hbar c}{mc^2} \times \begin{cases} \frac{1}{a_F} D, & \text{axial} \\ \frac{1}{d_p} D, & \text{planar} \end{cases} \quad (2)$$

$$D = \nu_0 = 2$$

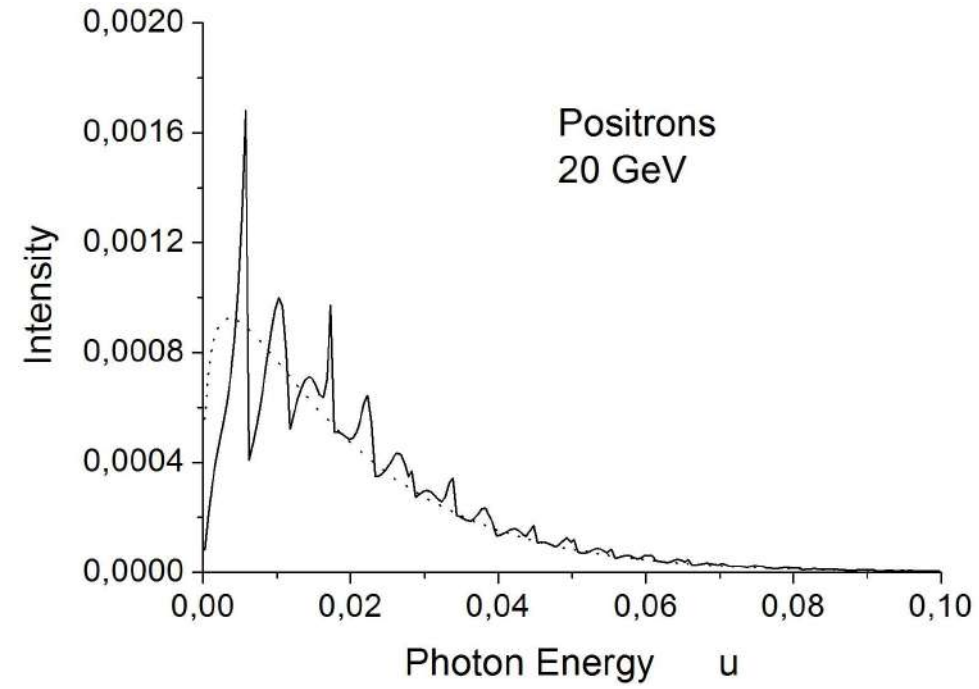
1 - $a = 1$, or 50 GeV

2 - $a = 6$, or 300 GeV



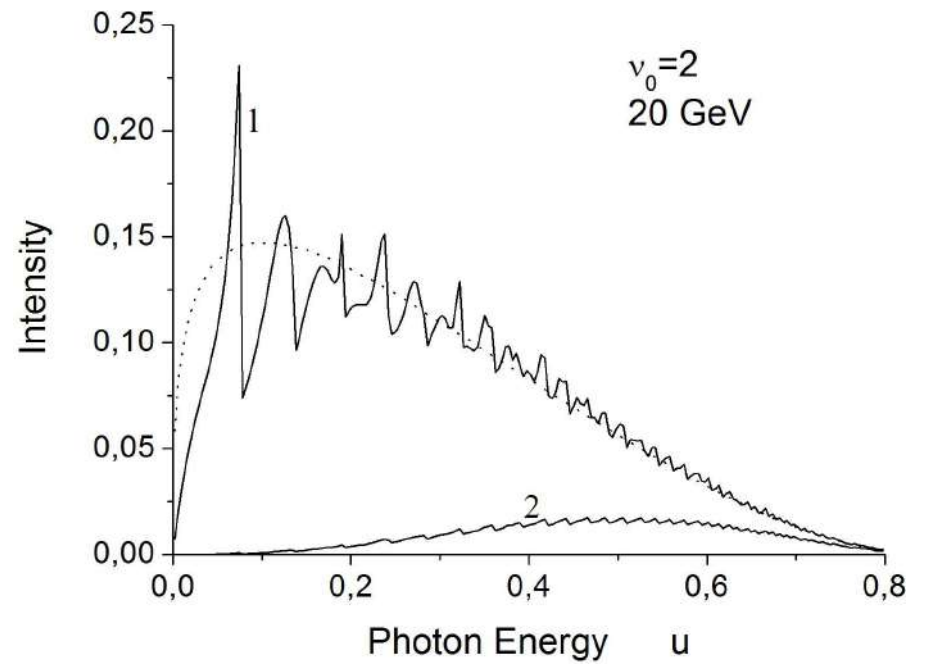
PLANAR

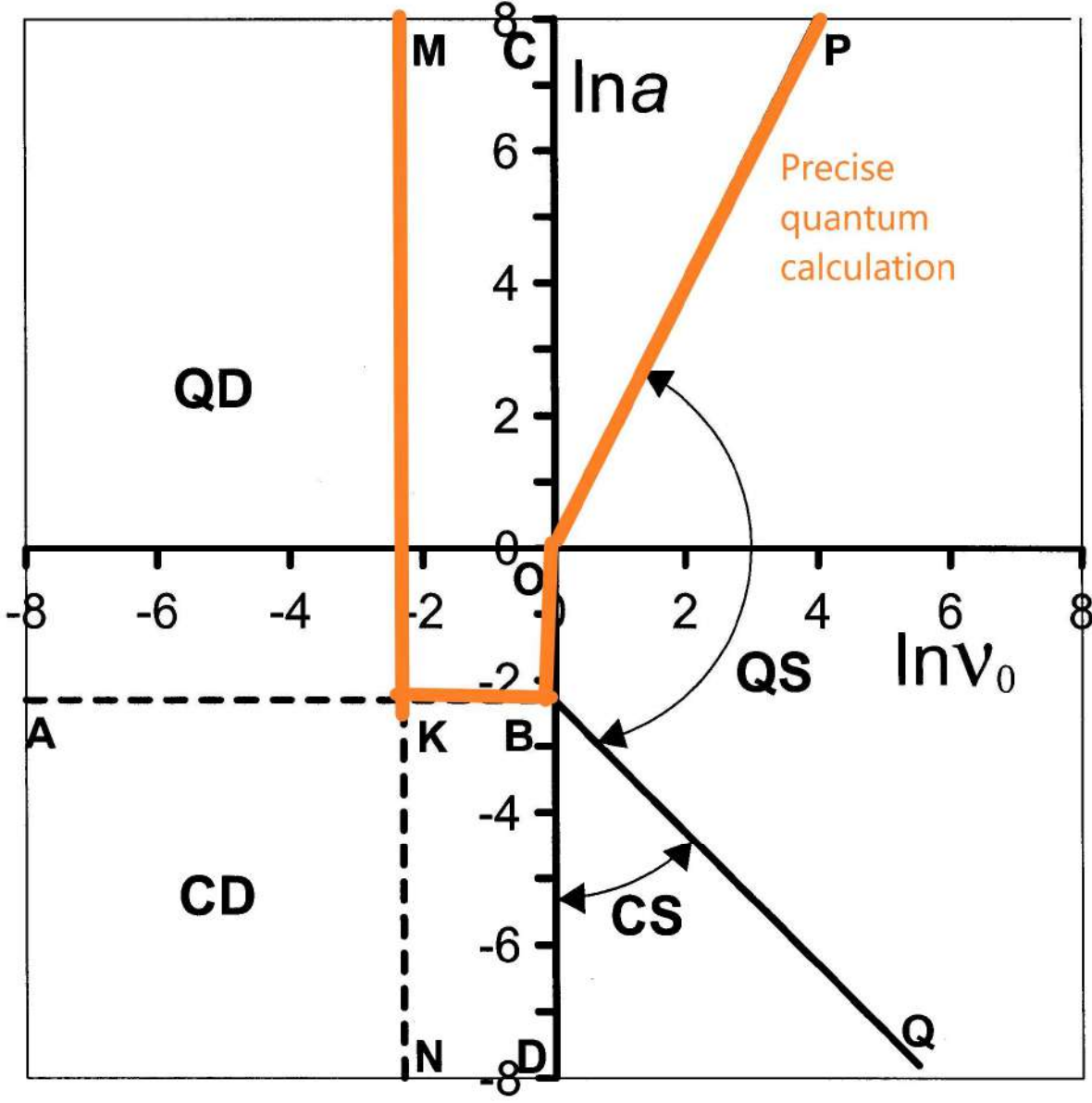
Positrons
20 GeV



LASER

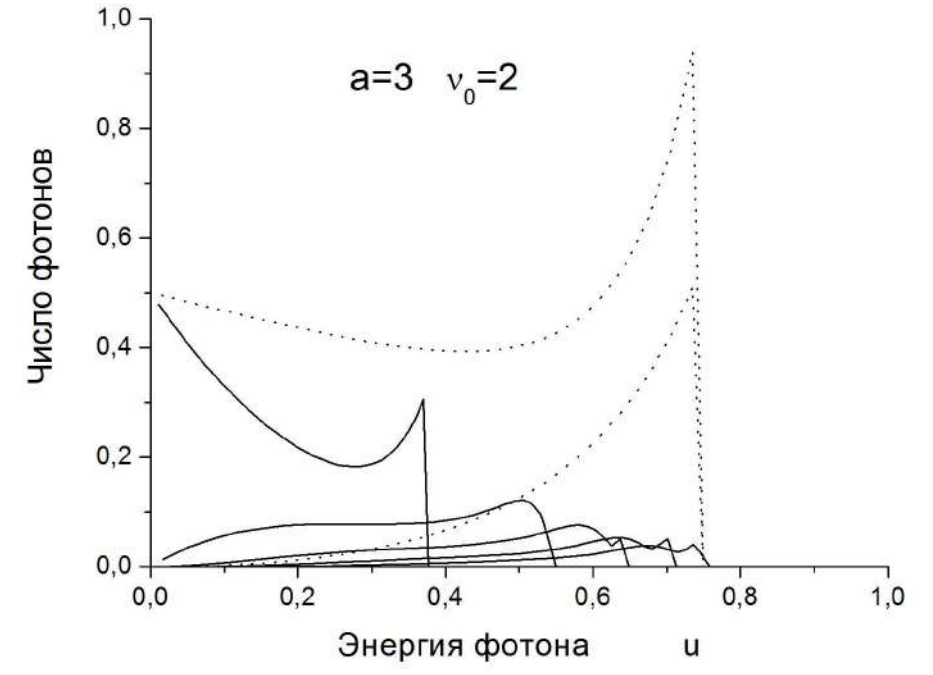
$\nu_0=2$
20 GeV

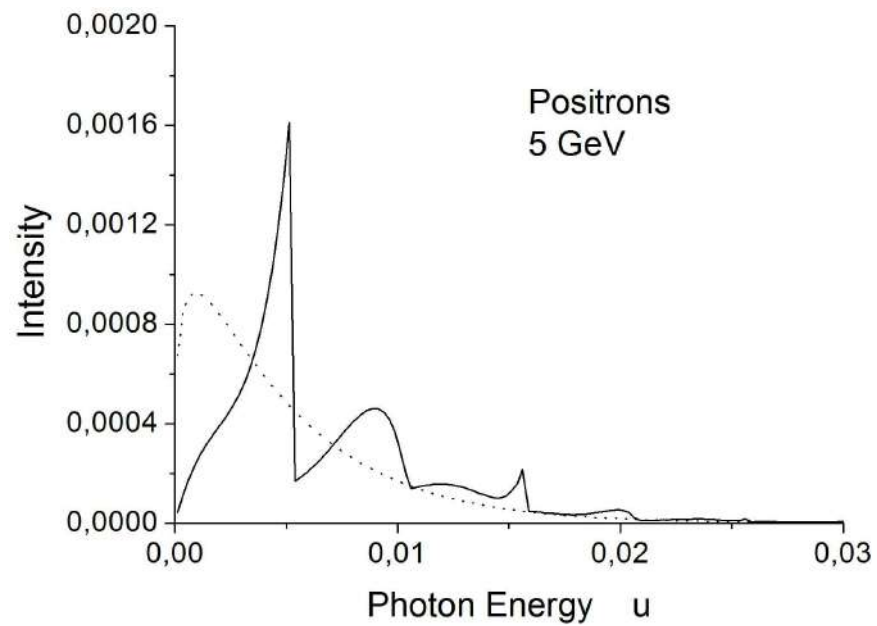
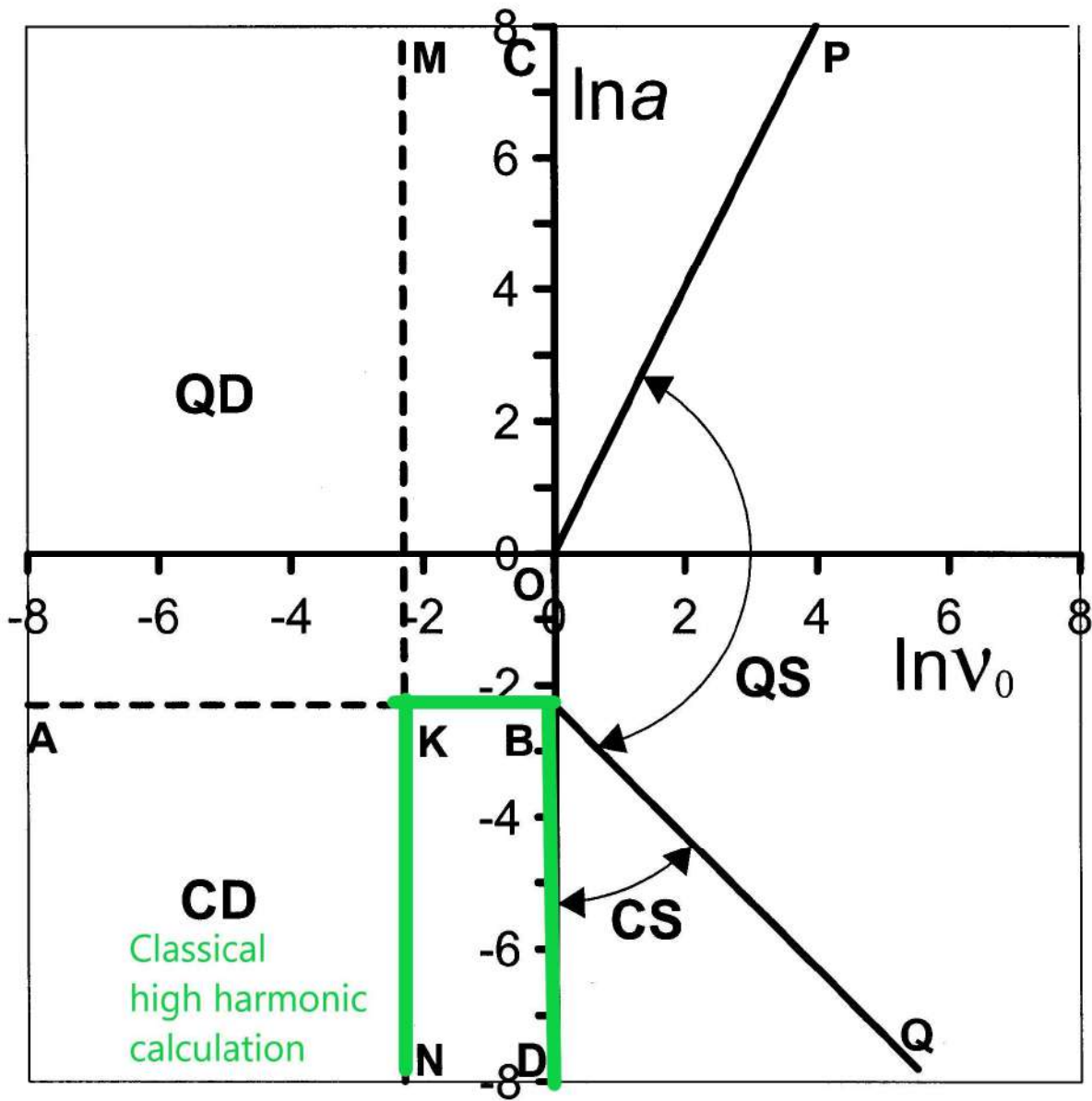


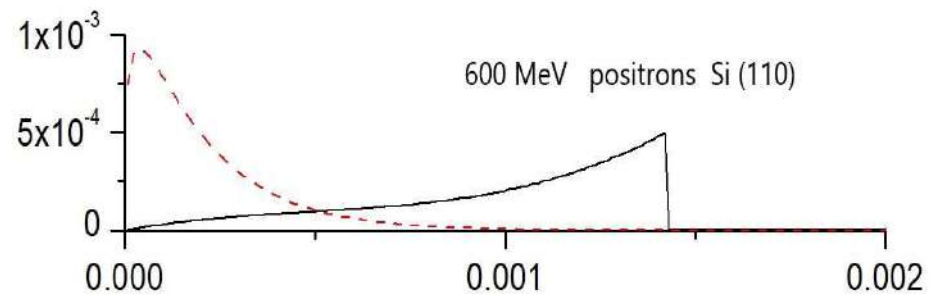
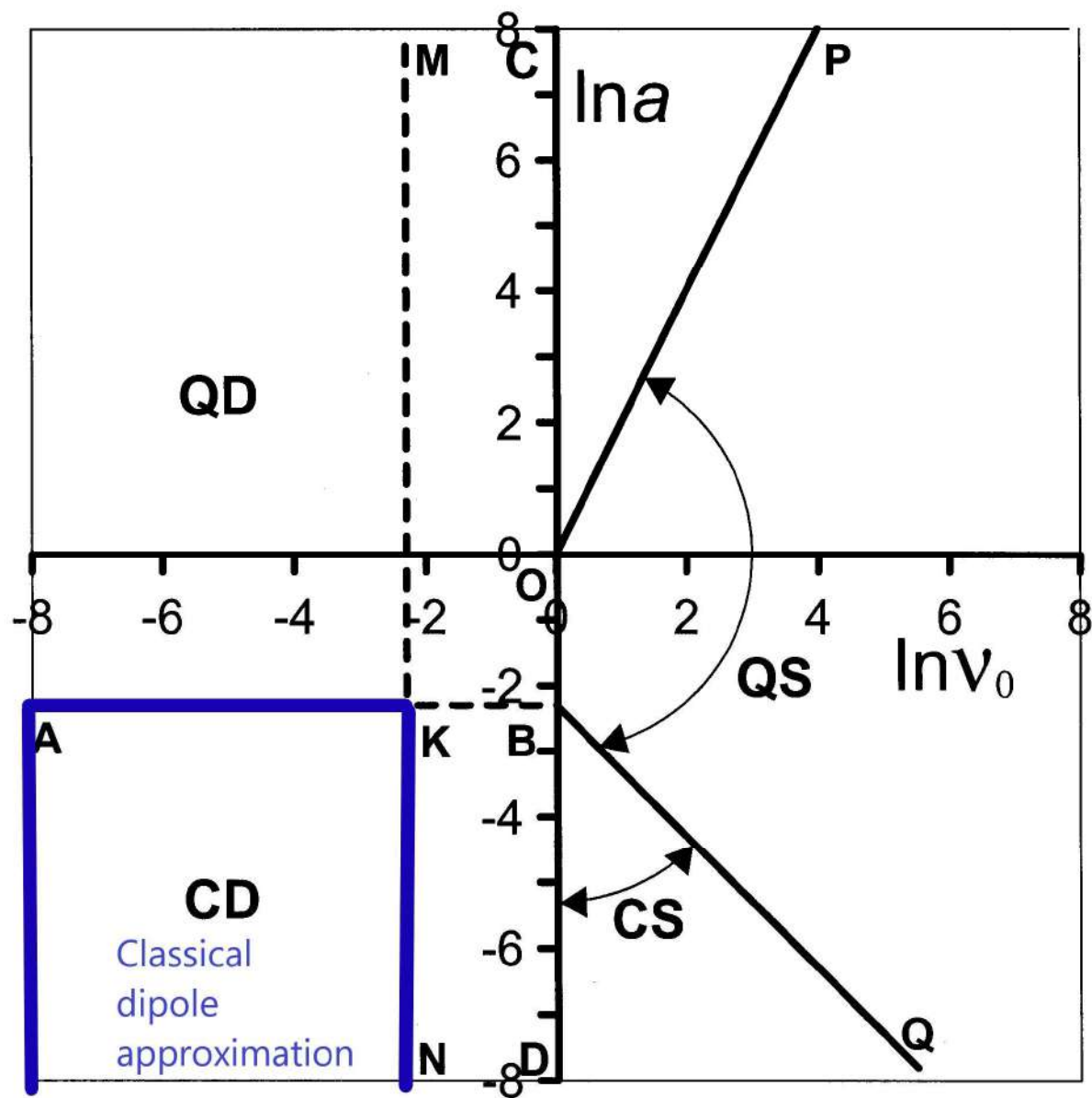


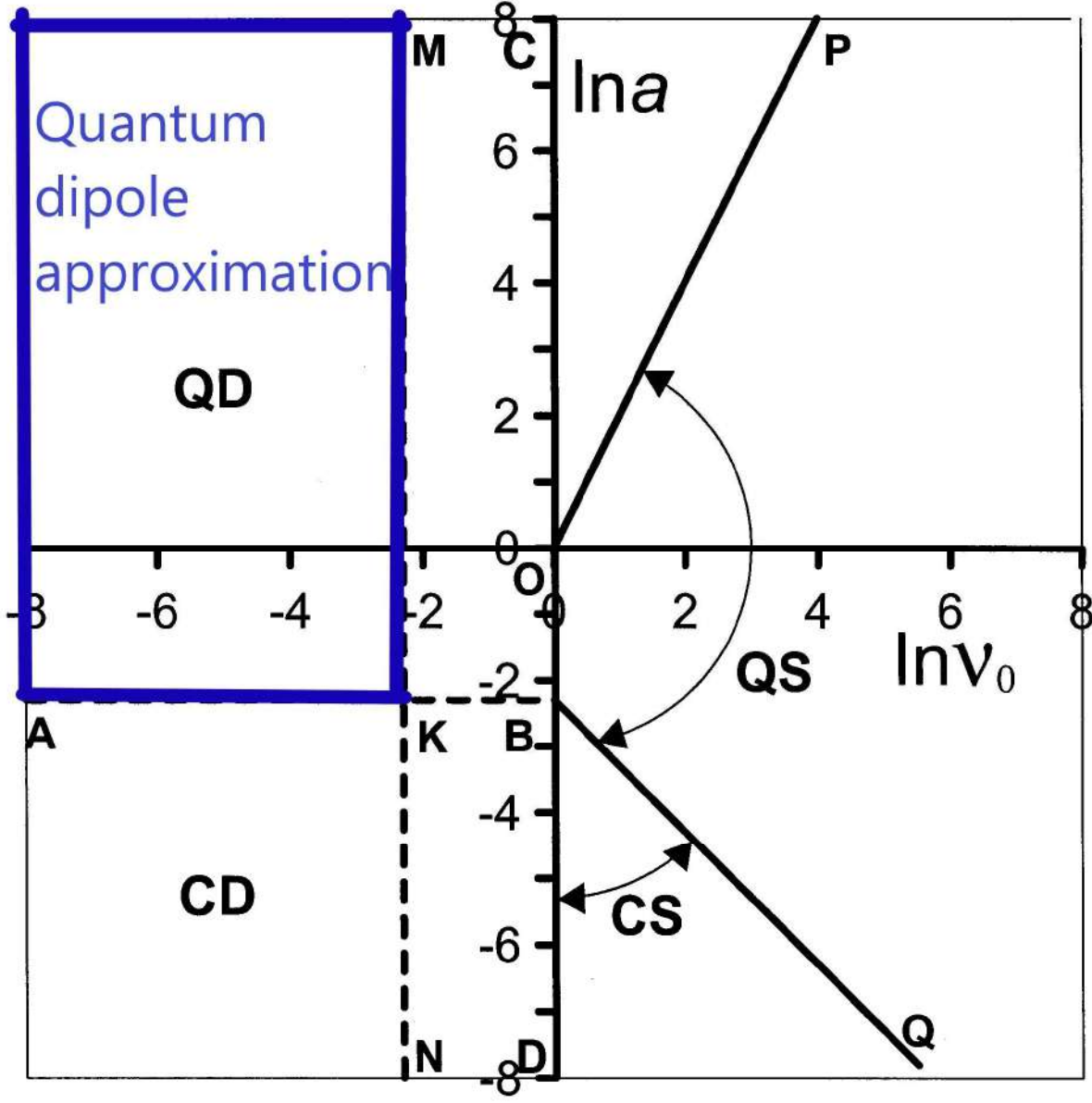
Not possible for channeling?

See previous figure 1

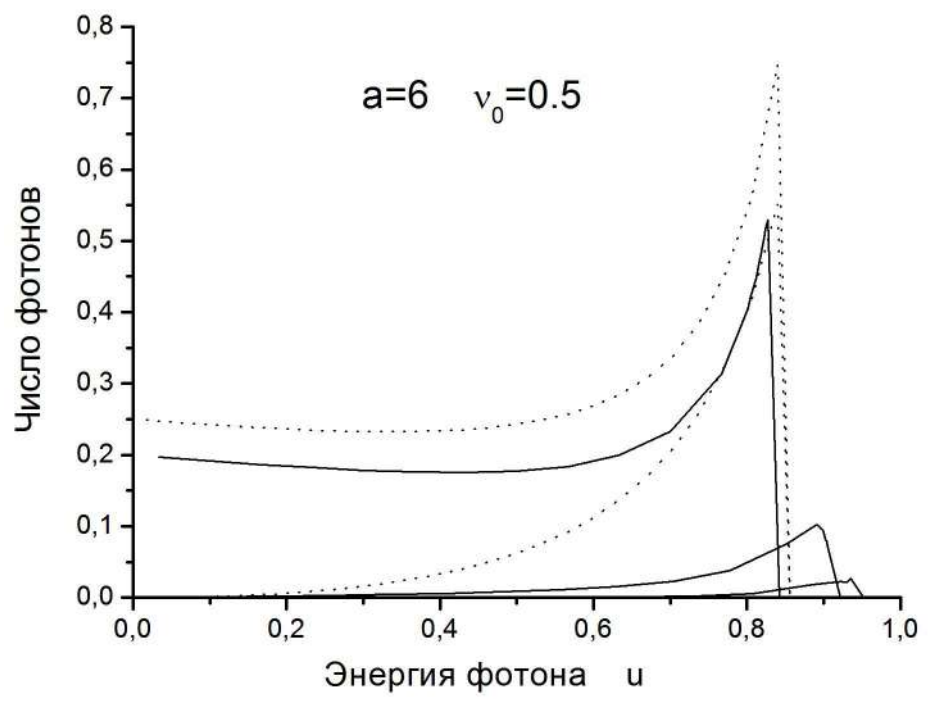


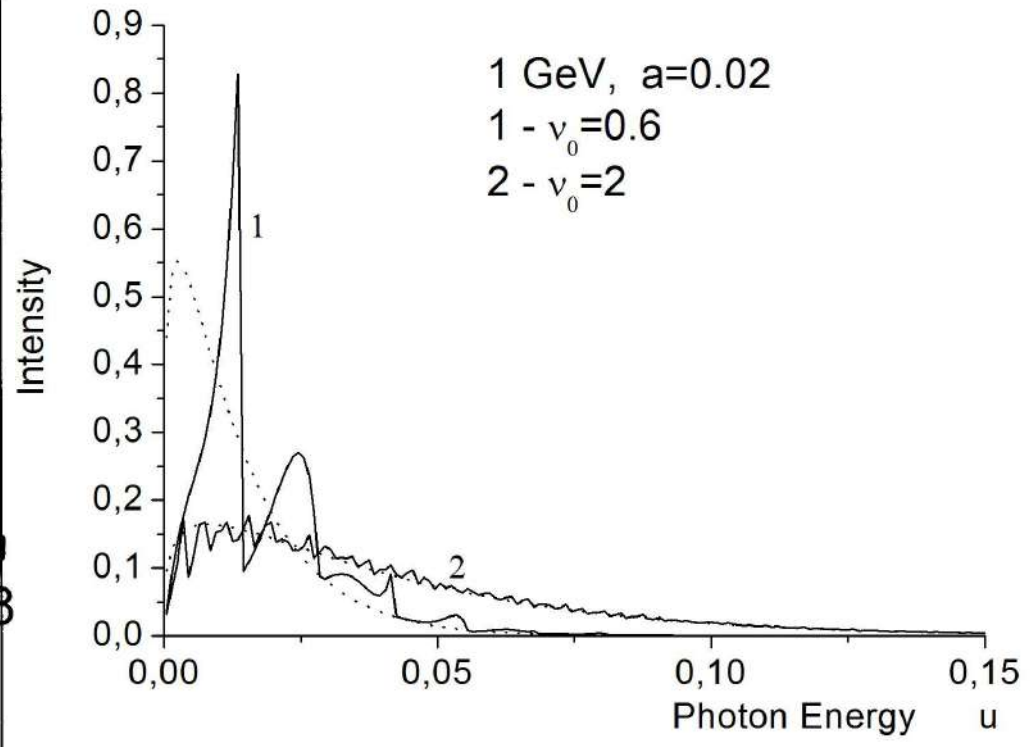
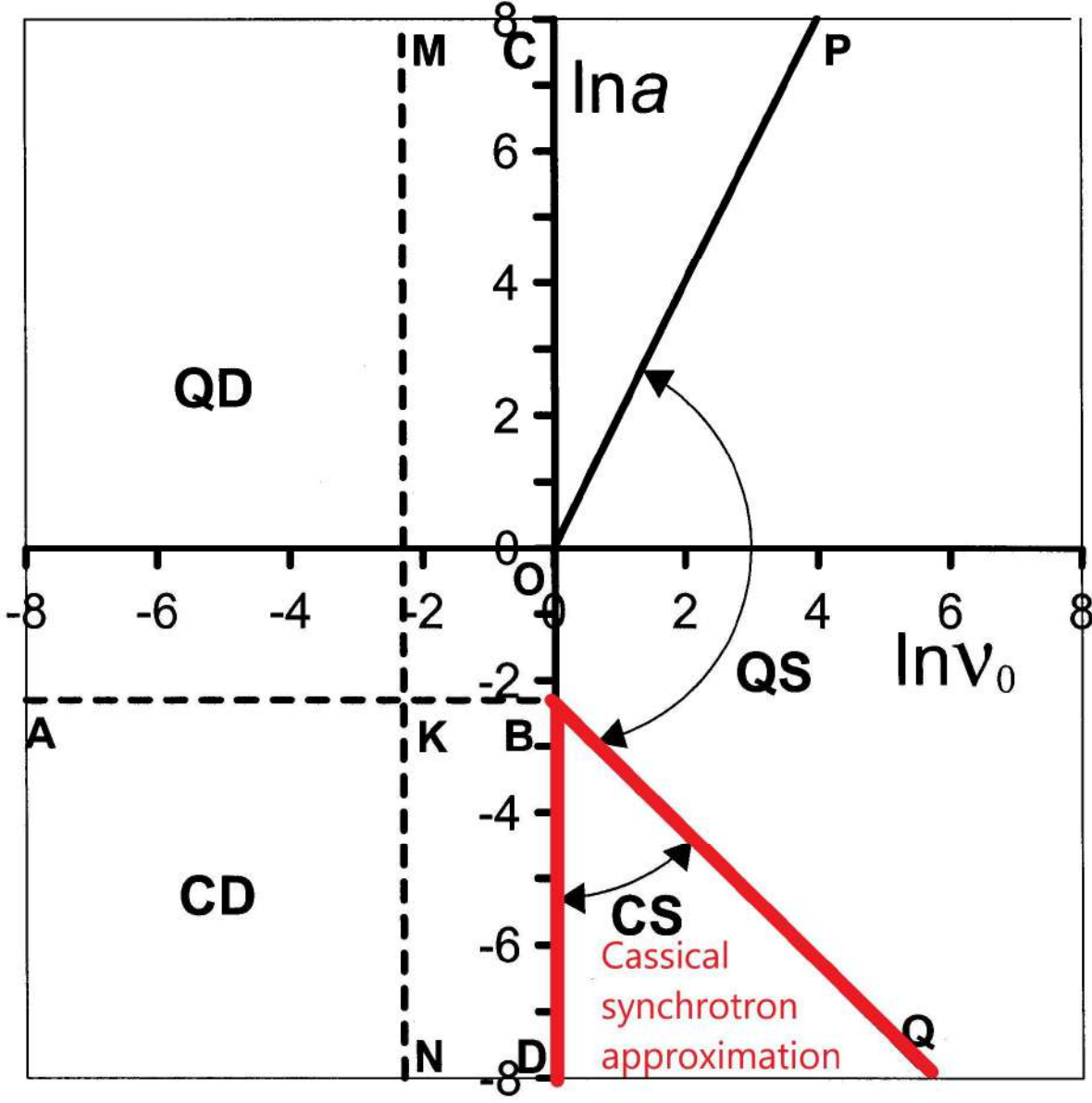




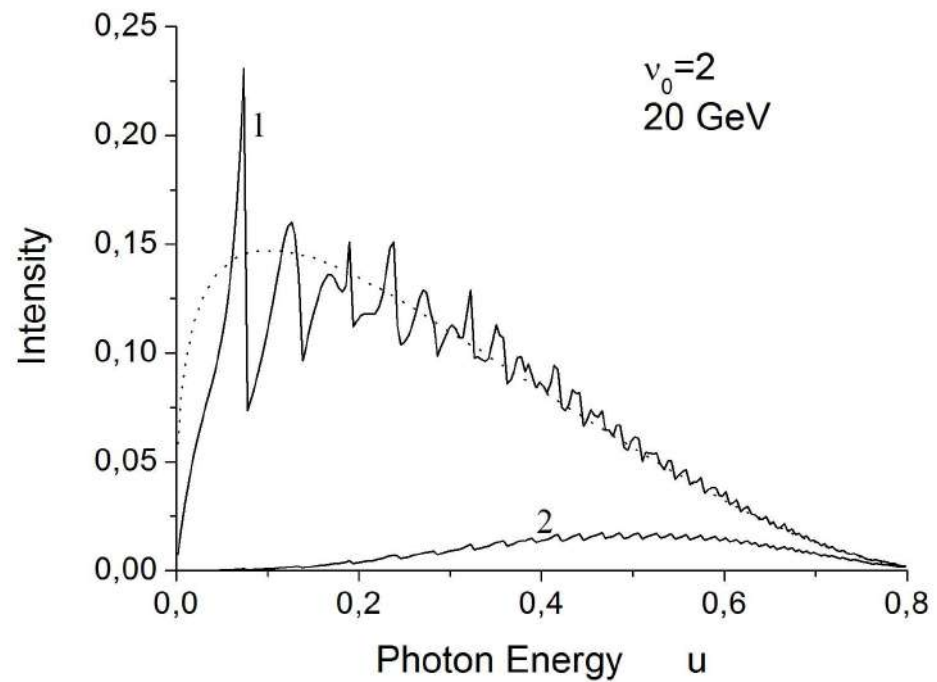
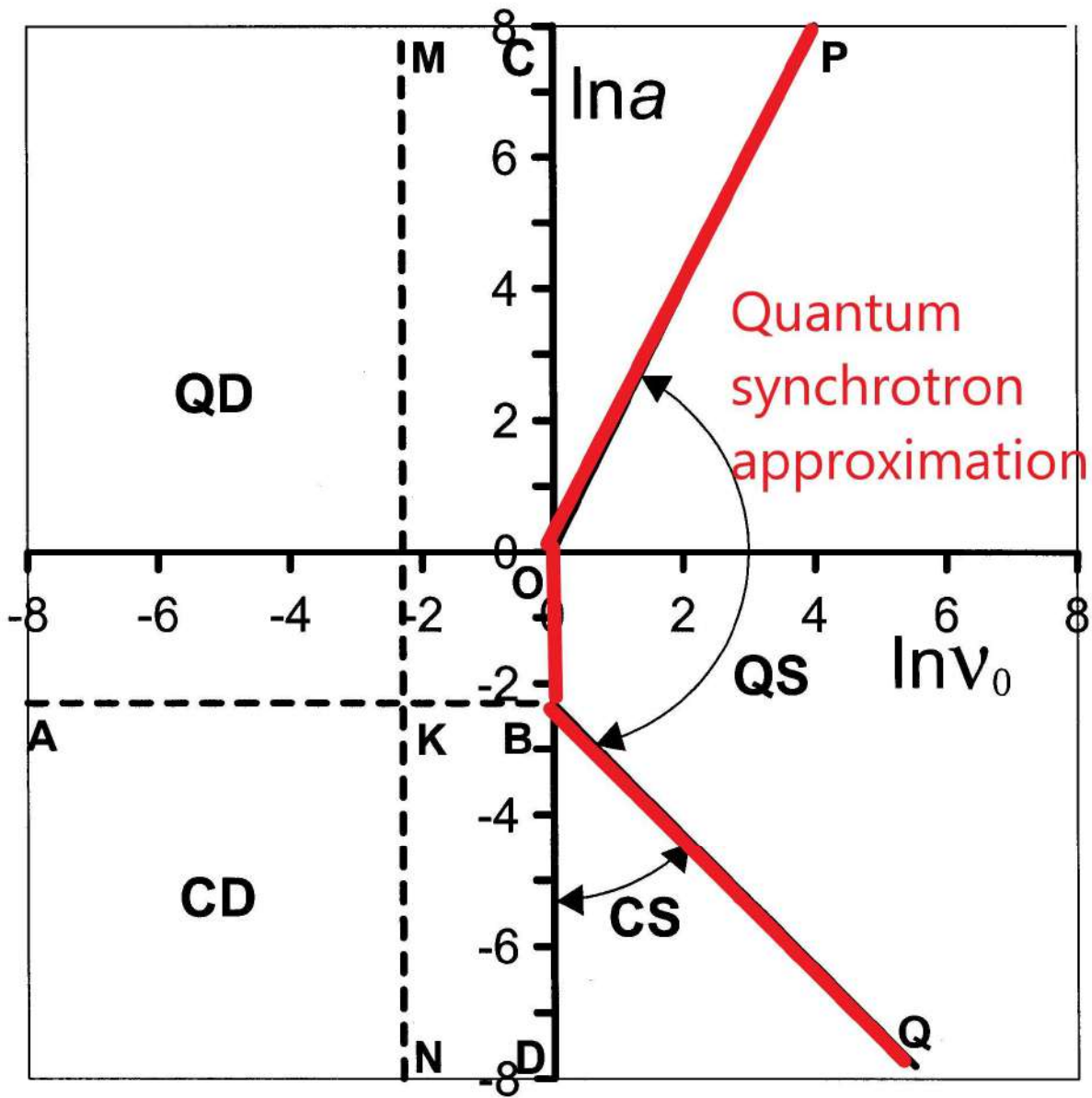


Not possible for channeling





See curve 2



Выводы

1. Параметры излучения для кристаллов и тераваттных лазеров почти совпадают для энергий до нескольких сот МэВ.
2. Спектры излучения для кристаллов и лазеров можно описать одним и тем же выражением, зависящим только от двух инвариантов a и D . В случае кристаллов область значений этих инвариантов сильно ограничена, так как они лежат на одной линии.
3. С ростом энергии пучка применимость синхротронного приближения для каналирования становится лучше, а для лазеров, наоборот.
4. При одной и той же энергии пучка излучение в поле лазеров смещено в область более жёстких частот.